

# KYUNG HEE UNIVERSITY

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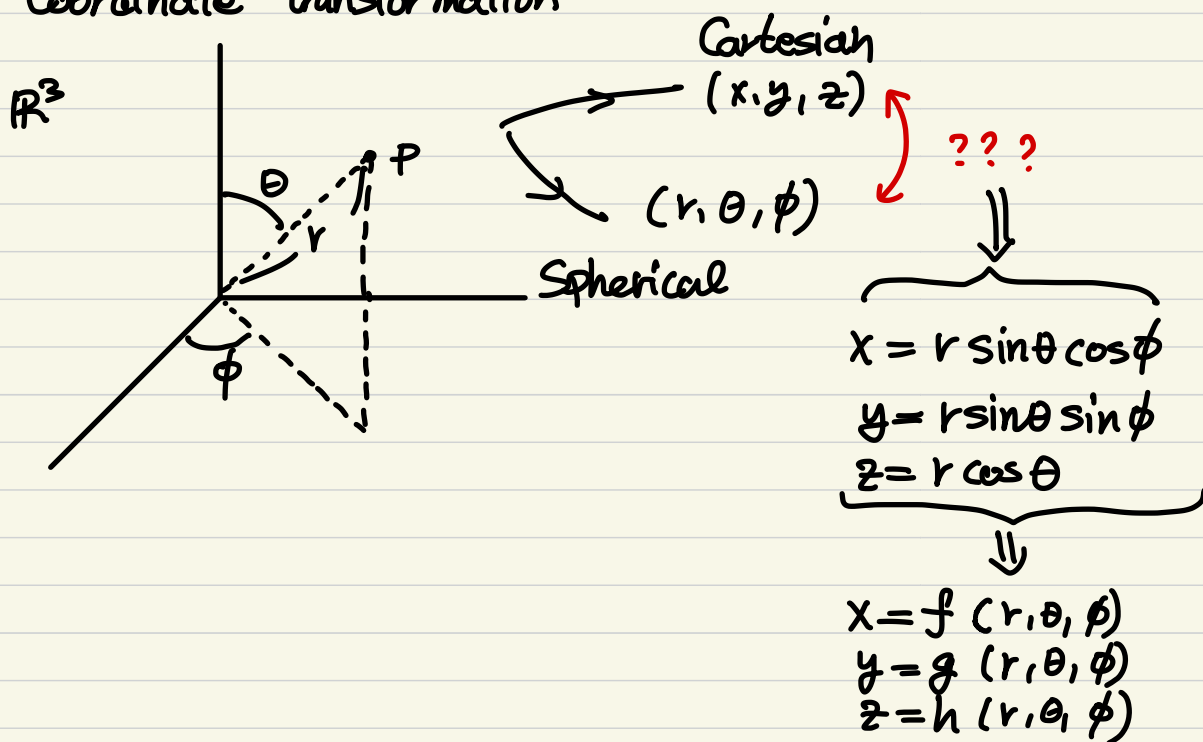


## Lectures on the Classical Field Theory



# I. Preliminaries

## I.1. Multilinear algebra - Coordinate transformation



In general, original coordinate system  $\{q_i = q_1, \dots, q_n\}$   
 changed coordinate system  $\{q'_i = q'_1, \dots, q'_n\}$

$q'_i = f'_i(q'_1, \dots, q'_n)$  : coordinate function

$$e'_j = \sum_{i=1}^n \frac{\partial q'_j}{\partial q_i} q_i = \sum_i \frac{\partial q'_j}{\partial q_i} q_i$$

Einstein's convention Repeated indices  $\Rightarrow$  summed index

$= A^i_j q_i$  : covariant transformation

$$\frac{\partial}{\partial q'_i} = \frac{\partial q'_j}{\partial q_i} \frac{\partial}{\partial q'_j} = A^i_j \frac{\partial}{\partial q'_j}$$

basis of vector  $e'_j$

$$dq_i = \frac{\partial q_i}{\partial q'_j} dq'_j = (A^{-1})^j_i dq'_j$$

basis of co-vector  $\bar{e}_j$  (dual)  $\hookrightarrow$  form

$\bar{e}_j \cdot e^i = \delta^i_j$

Contravariant transformation

$$\vec{A} \cdot \vec{B} = (A^j \bar{e}_j) \cdot (B_i e^i) = A^j B_i (\bar{e}_j \cdot e^i) = A^j B_i$$

$$\bar{e}_j = \langle j | \quad e^i = | i \rangle$$

$$A | i \rangle \leftrightarrow \langle j | A^{-1}$$

$\Rightarrow$  Under the transformation by  $A$ , the norm is not changed.

"Isometry"  $\leadsto$  group  $G$

ex)  $\mathbb{R}^3 \leadsto G = SO(3)$

$$e^i \rightarrow e'^i \quad \bar{e}_j \rightarrow \bar{e}'_j$$

### - Tensor

- Vector (representation of  $G$ )

$$V_{\bar{i}} \rightarrow A^{\bar{j}}_{\bar{i}} V_{\bar{j}}$$

- Covector

$$W^{\bar{i}} \rightarrow (A^{-1})^{\bar{i}}_{\bar{j}} W^{\bar{j}}$$

\* For  $\mathbb{R}^3$ , there is no "difference" between vector and co-vector due to the " $\bar{e} \cdot e = \delta$ ".

- $(p, q)$ -tensor  $(p+q)$ -rank tensor

$$T_{\bar{i}_1 \dots \bar{i}_p}^{\bar{j}_1 \dots \bar{j}_q} \rightarrow (A)^{k_1}_{\bar{i}_1} \dots (A)^{k_p}_{\bar{i}_p} (A^{-1})^{\bar{j}_1}_{l_1} \dots (A^{-1})^{\bar{j}_q}_{l_q} T_{\dots}$$

- \* (Anti-)Symmetric tensor

$$T_{\bar{i}_1 \dots \bar{i}_p \dots \bar{i}_q \dots \bar{i}_n} = \begin{matrix} \downarrow \text{symmetric} \\ \pm \\ \uparrow \text{anti-symmetric} \end{matrix} T_{\bar{i}_1 \dots \bar{i}_q \dots \bar{i}_p \dots \bar{i}_n}$$

$\begin{matrix} \uparrow p^{\text{th}} & & \uparrow p^{\text{th}} \\ \bar{i}_q & \dots & \bar{i}_p \end{matrix}$

ex) Matrix  $M$

$$M_{ij} = \pm M_{ji}$$

cf) We call an anti-symmetric  $(0, q)$ -tensor (rank  $q$  contravariant form) as  $q$ -form

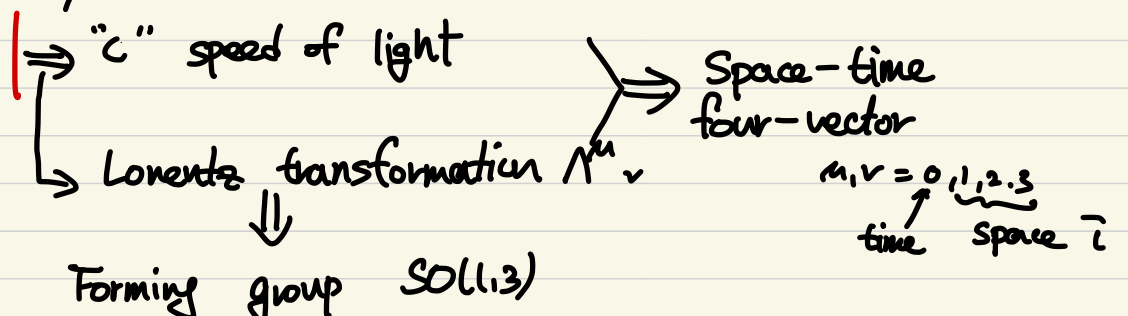
$$\omega^{\bar{i}_1 \dots \bar{i}_q} = -\omega^{\bar{i}_1 \dots \bar{i}_b \dots \bar{i}_a \dots \bar{i}_q}$$

## I.2. Special theory of Relativity

- Principle of Relativity (1905, Einstein)

"If a physics law hold in a coordinate system, it is also held to any other systems moving in uniform translation relatively to the reference one." ↗ frame moving with constant velocity

Maxwell theory  
Electrodynamics



- Minkowski metric

$$SO(0) \quad M^T \delta M = \mathbb{1}$$

$$SO(1,3) \quad M^T \underbrace{\begin{pmatrix} +1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix}}_{\text{Minkowski metric } \eta_{\mu\nu}} M = \mathbb{1}$$

$$\eta_{\mu\nu} = \text{diag}(+1, -1, -1, -1)$$

\* Inverse metric  $\eta^{\mu\nu}$   $\eta^{\mu\rho} \eta_{\rho\nu} = \delta^{\mu}_{\nu}$

\* Symmetric  $\eta_{\mu\nu} = \eta_{\nu\mu}$

\* Norm  $V \cdot W = V^{\mu} W^{\nu} \eta_{\mu\nu} = V^0 W^0 - V^1 W^1 - V^2 W^2 - V^3 W^3$

\* One-to-one map between vector  $\leftrightarrow$  covector

$$V_{\mu} = \eta_{\mu\nu} V^{\nu} \quad W^{\mu} = \eta^{\mu\nu} W_{\nu}$$

(i.e.  $V_0 = V^0$ ,  $V_i = -V^i$ , ...)

$$T_{\mu\nu} = T_{\mu}^{\rho} \eta_{\rho\nu}$$

• Four-vector

$$X^{\mu} = (x^0, x^1, x^2, x^3)$$

$$X^2 = X^{\mu} X_{\mu} = X^{\mu} X^{\nu} \eta_{\mu\nu} = (x^0)^2 - (x^1)^2 - (x^2)^2 - (x^3)^2$$

ex) position  $(ct, x, y, z) = x^{\mu}$

$$X^2 = c^2 t^2 - x^2 - y^2 - z^2$$

energy-momentum  $(cE, p_x, p_y, p_z) = p^{\mu}$

$$c^2 E^2 - (p_x)^2 - (p_y)^2 - (p_z)^2 = m^2 c^4$$

Under Lorentz transformation

$$\begin{cases} V'^{\mu} = \Lambda^{\mu}_{\nu} V^{\nu} \\ W'_{\mu} = (\Lambda^{-1})_{\mu}^{\nu} W_{\nu} \end{cases}$$

$$\eta'_{\mu\nu} = (\Lambda^{-1})_{\mu}^{\rho} (\Lambda^{-1})_{\nu}^{\sigma} \eta_{\rho\sigma}$$

⋮



## II. Klein-Gordon Equation

II.1. Klein-Gordon Equation  
Special relativity (1905)

Schrödinger equation (1925)

⇒ Klein-Gordon equation (1926)  
⇓

$$-i \frac{\partial}{\partial t} \psi = H \psi$$

Dirac equation (1928)

• First trial  $\rightarrow E$

$$-i \frac{\partial}{\partial t} \psi = H \psi = \pm \sqrt{-\vec{v}^2 + m^2} \psi$$

two weaknesses

① negative sign

② Non-locality

⇒ square both side  $\rightarrow$  Klein-Gordon equation

$$-\frac{\partial^2}{\partial t^2} \psi = -\vec{v}^2 + m^2 \psi$$

$$\left( \frac{\partial^2}{\partial t^2} - \vec{v}^2 \right) \psi + m^2 \psi = \left( \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2} \right) \psi + m^2 \psi = 0$$

$$\square \equiv \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2} = \eta_{\mu\nu} \frac{\partial}{\partial x^\mu} \frac{\partial}{\partial x^\nu}$$

: d'Alembertian

$$(\square + m^2) \phi = 0 \quad ; \text{ Klein-Gordon equation}$$

$\rightarrow$  mass of field  $\phi$

- Fourier decomposition

$$\phi(x) = \int_{\mathbb{R}^3} \frac{d^3\vec{k}}{(2\pi)^{3/2}} \varphi_{\vec{k}}(t) e^{-i\vec{k}\cdot\vec{x}}$$

Reality condition:  $\phi^*(x) = \phi(x)$

$$\phi^*(x) = \int \frac{d^3\vec{k}}{(2\pi)^{3/2}} \varphi_{\vec{k}}^*(t) e^{+i\vec{k}\cdot\vec{x}} = \phi(x)$$

⇒  $\varphi_{\vec{k}}(t) = \varphi_{-\vec{k}}(t)$  : Reality condition

To manifest Lorentz covariance

$$\varphi_{\vec{k}} = \frac{1}{\sqrt{2\omega_{\vec{k}}}} (a_{\vec{k}} e^{+i\omega_{\vec{k}}t} + b_{\vec{k}} e^{-i\omega_{\vec{k}}t}) \quad \omega_{\vec{k}}^2 = k^2 + m^2$$

Reality condition ⇒  $b_{-\vec{k}} = a_{\vec{k}}^*$

$$\Rightarrow \phi(x) = \int \frac{d^3\vec{k}}{(2\pi)^{3/2}} \frac{1}{\sqrt{2\omega_{\vec{k}}}} \left( a_{\vec{k}} e^{i\vec{k}\cdot\vec{x}} + a_{\vec{k}}^* e^{-i\vec{k}\cdot\vec{x}} \right)$$

Quantization  $\left\{ \begin{array}{l} \text{negative} \\ \text{frequency} \end{array} \right. \rightarrow$  anti-particle / annihilation  
 $\left\{ \begin{array}{l} \text{positive} \\ \text{frequency} \end{array} \right. \rightarrow$  particle / creation

• Solving Klein-Gordon

$$(\square + m^2) \phi(x) = 0$$

$$\Rightarrow \frac{\partial^2}{\partial t^2} \varphi_{\vec{k}}(t) + (+\vec{k}^2 + m^2) \varphi_{\vec{k}}(t) = 0$$

$\Rightarrow$  Harmonic oscillator

- Note
- We can deal with a (classical) field as a collection of infinite harmonic oscillators.
  - After quantization,  $a$  and  $a^*$  are promoted to the harmonic oscillator creation / annihilation operators  $a$  and  $a^\dagger$ .

## II.2. Scalar Lagrangian

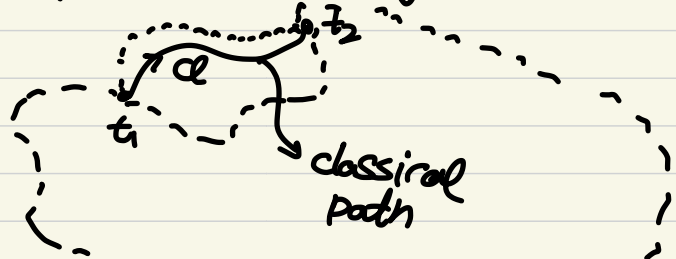
- Recall: Lagrangian mechanics

Hamilton's principle  $\Rightarrow$  Classical path = extremising action functional

$$S[x] = \int_{t_1}^{t_2} dt L(x, \dot{x})$$

$$\delta S \stackrel{!}{=} 0$$

$$\Rightarrow S[x + \delta x] - S[x] = 0$$



$$\begin{aligned} S[x + \delta x] &= \int_{t_1}^{t_2} dt L(x + \delta x, \dot{x} + \delta \dot{x}) \\ &= \int_{t_1}^{t_2} dt \left( \frac{\partial L}{\partial x} \delta x + \frac{\partial L}{\partial \dot{x}} \delta \dot{x} \right) + \cancel{\mathcal{O}(\delta^2)} + S[x] \\ &= \int_{t_1}^{t_2} dt \left( \frac{\partial L}{\partial x} - \frac{d}{dt} \frac{\partial L}{\partial \dot{x}} \right) \delta x + \left( \frac{\partial L}{\partial \dot{x}} \right) \Big|_{t_1}^{t_2} + S[x] \end{aligned}$$

$$\delta S = 0, \forall \delta x \Rightarrow \boxed{\frac{\partial L}{\partial x^i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{x}^i} = 0} \quad \text{Euler-Lagrange equation}$$

$$\frac{\partial L}{\partial \dot{x}^i} = p_i : \text{canonical momenta}$$

- Field equation from the Lagrangian

$$S[\phi] = \int dt L(\phi, \partial \phi) \quad \partial_\mu = \frac{\partial}{\partial x^\mu}$$

not Lorentz covariant form

Lagrangian density

$$= \int d^4x \mathcal{L}(\phi, \partial \phi) \quad \mathcal{L}(\phi, \partial \phi) = \int d^3\vec{x} L(\phi, \partial \phi)$$

$$\delta S = 0 \Rightarrow S[\phi + \delta \phi] - S[\phi] = 0$$

$$= \int d^4x [L(\phi + \delta \phi, \partial \phi + \delta(\partial \phi)) - L(\phi, \partial \phi)]$$

$$\begin{aligned}
&= \int d^4x \left( \frac{\partial \mathcal{L}}{\partial \phi} \delta \phi + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \delta (\partial_\mu \phi) \right) \\
&= \int d^4x \left( \frac{\partial \mathcal{L}}{\partial \phi} \delta \phi - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \delta \phi \right) - \left[ \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right]_{\text{bdy}} \\
&= \int d^4x \left( \frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right) \delta \phi = 0, \forall \delta \phi \\
\Rightarrow \quad &\boxed{\frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} = 0} \quad \text{Euler-Lagrange equation}
\end{aligned}$$

- Scalar Lagrangian

$$\begin{aligned}
\mathcal{L}(\phi, \partial \phi) &= -\frac{1}{2} (\partial_\mu \phi) (\partial^\mu \phi) + \frac{1}{2} m^2 \phi^2 \\
&= \frac{1}{2} \phi (\square + m^2) \phi + (\text{boundary})
\end{aligned}$$

$$\frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} = (m^2 + \square) \phi = 0 \Rightarrow \text{Klein-Gordon equation}$$

### II.3. Complex Scalar theory

$\phi_1, \phi_2$  : Real scalar field

$\phi = \phi_1 + i\phi_2 \Rightarrow$  Complex scalar field

Condition: Lagrangian should be real and Lorentz invariant

$$\mathcal{L}(\phi, \phi^*) = (\partial_\mu \phi) (\partial^\mu \phi^*) + m^2 |\phi|^2$$

$$\begin{cases} (\square + m^2) \phi = 0 \\ (\square + m^2) \phi^* = 0 \end{cases}$$

Note After quantization,  $\phi$  can be regarded as a particle and  $\phi^*$  can be regarded as an anti-particle.

particle annihilation  
anti-particle creation

particle creation  
anti-particle annihilation

$$\begin{cases} \phi = \int \frac{d^3k}{\sqrt{2\omega_k}} \frac{1}{(2\pi)^{3/2}} \left( \overline{a_k} e^{ik \cdot x} + a_k^* e^{-ik \cdot x} \right) \\ \phi^* = \int \frac{d^3k}{\sqrt{2\omega_k}} \frac{1}{(2\pi)^{3/2}} \left( b_k e^{ik \cdot x} + \overline{b_k^*} e^{-ik \cdot x} \right) \end{cases}$$

particle creation  
anti-particle annihilation

particle annihilation  
anti-particle creation

- Gauge theory : First-look

Recall In Maxwell theory

$$A_\mu = (\phi, \vec{A}) \rightarrow A_\mu + \partial_\mu \alpha = \left( \phi - \frac{\partial \alpha}{\partial t}, \vec{A} + \vec{\nabla} \alpha \right)$$

[ Maxwell equation  
Electric/Magnetic field ] has not changed under this transf.

In QM

$$|\psi\rangle \rightarrow e^{i\alpha} |\psi\rangle$$

↳ [ Schrödinger equation  
Probability ] does not change

⇒ Under this transformation, physical quantities are invariant  
: Gauge transformation

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$$\phi \rightarrow e^{i\alpha} \phi, \quad \phi^* \rightarrow e^{-i\alpha} \phi^* \quad \therefore \text{Global } U(1) \text{ transf.}$$

$$\left[ \begin{array}{l} \mathcal{L} = (\partial_\mu \phi) (\partial^\mu \phi^*) + m^2 |\phi|^2 \\ (\square + m^2) \phi = 0 \end{array} \right] \Rightarrow \text{Unchanged under transformation.}$$

⇒ Complex scalar theory is a first example of gauge theory

### III. Noether Theorem

#### III.1. The Law of Conservation

$$\bar{j}^\mu = j^\mu(\phi, \partial\phi) \quad ; \text{ conserved current}$$

• Conservation law

$$\partial_\mu j^\mu = 0 \quad \text{when } (\square - m^2)\phi = 0$$

$$\text{Let } j^\mu = (\rho, \vec{J})$$

$$\partial_\mu j^\mu = \frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{J} = 0 \quad ; \text{ continuity equation}$$

• Charge

$$Q_V = \int_V d^3x \rho(x)$$

$$\frac{d}{dt} Q_V = \int_V d^3x \frac{d}{dt} \rho(x)$$

$$= - \int_V d^3x \vec{\nabla} \cdot \vec{J} = - \int_{\partial V} d^2\vec{S} \cdot \vec{J}$$

$\Rightarrow$  If  $Q_V$  is a conserved charge, then

$Q_V$  is never neither created or annihilated.

$\Rightarrow$  We can regard the "total charge" of the region as a constant of motion.

E.g.)  $\bar{j}^\mu$  : electric current  $\xleftrightarrow{\text{Noether than.}}$   $U(1)$   
energy - momentum  $\xleftrightarrow{\text{Noether than.}}$   $R^{1,3}$   
angular momentum  $\xleftrightarrow{\text{Noether than.}}$   $SO(1,3)$   
:

#### III.2. Transformation of Fields

- Transformation

• Case 1: vector rotation

$$V = \begin{pmatrix} a \\ b \end{pmatrix} \xrightarrow{\text{2d rotation}} V' = \begin{pmatrix} a' \\ b' \end{pmatrix} = \overbrace{\begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix}}^R \begin{pmatrix} a \\ b \end{pmatrix} = \begin{pmatrix} a\cos\theta - b\sin\theta \\ a\sin\theta + b\cos\theta \end{pmatrix}$$

$$\Rightarrow V'^i = R(\theta)^i_j V^j \quad ; \text{ (SO(2)) transformation}$$

• Case 2: Phase shift of wave function

$$\psi \longrightarrow \psi' = e^{i\theta} \psi$$

$$\text{In general, } \phi \longrightarrow \phi' = f(\theta, \tau) \phi$$

[ Global  $\theta_i = \theta$  ] [ Continuous  $f(\theta_i)$ : cont.  $\rightarrow$  Lie group ]  
 [ Local  $\theta_i = \theta(x)$  ] [ Discrete  $f(\theta_i)$ : disc.  $\rightarrow$  Discrete group ]

• Infinitesimal transformation

$$\frac{\delta \phi}{\delta \theta_i} \equiv \delta_i \phi = \left. \frac{\partial \phi'}{\partial \theta_i} \right|_{\theta_i=0}$$

$$\delta_i \phi = \underbrace{\left( \frac{\partial f}{\partial \theta_i} \right) \Big|_{\theta_i=0}}_{\equiv \alpha_i} \phi \equiv \alpha_i \phi \cdot \text{Linearised equation}$$

$\equiv \alpha_i \rightarrow$  Lie algebra

ex) space-time translation

$$\phi'(x^\mu) = \phi(x^\mu + \Delta^\mu) \Rightarrow \delta_{\Delta^\mu} \phi = \partial_\mu \phi$$

ex) scaling transformation

$$\phi'(x) = e^\lambda \phi \Rightarrow \delta_\lambda \phi = \phi$$

ex) 3d rotation of vector

$$\vec{V}' = R_z(\alpha) R_y(\beta) R_x(\gamma) \vec{V}$$

$$R_z(\alpha) = \begin{pmatrix} \cos \alpha & -\sin \alpha & \\ \sin \alpha & \cos \alpha & \\ & & 1 \end{pmatrix} \quad R_y(\beta) = \begin{pmatrix} \cos \beta & 0 & \sin \beta \\ & 1 & \\ -\sin \beta & & \cos \beta \end{pmatrix}$$

$$R_x(\gamma) = \begin{pmatrix} 1 & & \\ & \cos \gamma & -\sin \gamma \\ & \sin \gamma & \cos \gamma \end{pmatrix}$$

HW. Let  $\delta_\alpha \vec{V} = T_z V$ ,  $\delta_\beta \vec{V} = T_y V$ ,  $\delta_\gamma \vec{V} = T_x V$   
 Then what is the commutator relation between  $\delta T_x, T_y, T_z$ ?

### III.3. The Noether's (First) theorem

Consider  $\mathcal{L} = \mathcal{L}(\phi, \partial \phi)$ , Symmetry transformation  $\delta_i \phi$

$$\begin{aligned} \delta_i \mathcal{L} &\stackrel{!}{=} \partial_\nu W^\nu = \frac{\partial}{\partial \theta_i} \mathcal{L}(\phi', \partial \phi') \Big|_{\theta_i=0} \\ &= \frac{\partial \mathcal{L}}{\partial \phi'} \frac{\partial \phi'}{\partial \theta_i} + \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi')} \frac{\partial (\partial_\nu \phi')}{\partial \theta_i} \Big|_{\theta_i=0} \\ &= \frac{\partial \mathcal{L}}{\partial \phi} \delta_i \phi - \partial_\nu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \delta_i \phi \right) + \partial_\nu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \delta_i \phi \right) \\ &= \left( \frac{\partial \mathcal{L}}{\partial \phi} - \partial_\nu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \right) \right) \delta_i \phi + \partial_\nu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \delta_i \phi \right) \end{aligned}$$

when on-shell  
(eom is held)

$$\Rightarrow \partial_\mu \overbrace{(j_i^\mu - W^\mu)} = - \left( \frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right) \stackrel{!}{=} 0$$

$$\Rightarrow \boxed{j_i^\mu = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \delta_i \phi - W^\mu} \quad \text{Noether's first theorem.}$$

Noether current  
corresponding to the symmetry (generator)  $\Theta_i$

$$Q_i = \int d^3\vec{x} J_i^0 \quad : \text{Noether charge}$$

### III.4. Examples of Noether current

- Translation symmetry

$$\delta_\epsilon \phi = \partial_\epsilon \phi$$

$$\mathcal{L} = -\frac{1}{2} (\partial_\nu \phi) (\partial^\nu \phi) + \frac{1}{2} m^2 \phi^2$$

$$\begin{aligned} \delta_\mu \mathcal{L} &= -\frac{1}{2} (\partial_\nu \partial^\mu \phi) (\partial^\nu \phi) - \frac{1}{2} (\partial_\nu \phi) (\partial^\nu \partial^\mu \phi) + \frac{1}{2} m^2 ((\partial^\mu \phi) \phi + \phi \partial^\mu \phi) \\ &= -\frac{1}{2} \partial^\mu [(\partial_\nu \phi) (\partial^\nu \phi)] + \frac{1}{2} m^2 \partial^\mu (\phi^2) \\ &= \partial^\mu \eta_{\mu\epsilon} \mathcal{L} \end{aligned}$$

$$\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} = -\delta_\nu^\mu \partial^\nu \phi = -\partial^\mu \phi$$

$$\Rightarrow J^\mu_\epsilon = (-\partial^\mu \phi) (\partial_\epsilon \phi) - \delta^{\mu\epsilon} \eta_{\nu\rho} \mathcal{L}$$

$$J^{\mu\nu} = \eta^{\nu\rho} J^\mu_\epsilon = -\partial^\mu \phi \partial^\nu \phi - \delta^{\mu\nu} \mathcal{L} \equiv T^{\mu\nu} \quad : \text{energy-momentum tensor}$$

$$\begin{aligned} \partial_\mu J^\mu_\epsilon &= -\square \phi (\partial_\epsilon \phi) - \partial^\mu \phi (\partial_\mu \partial_\epsilon \phi) - \partial_\epsilon \mathcal{L} \\ &= m^2 \phi (\partial_\epsilon \phi) - \partial^\mu \phi (\partial_\mu \partial_\epsilon \phi) \\ &\quad + (\partial_\epsilon \partial_\nu \phi) (\partial^\nu \phi) - m^2 (\phi \partial_\epsilon \phi) = 0 \end{aligned}$$

- Lorentz symmetry

$\Rightarrow$  Preserve Minkowski metric

$$\Lambda^\mu_\epsilon \Lambda^\nu_\sigma \eta_{\mu\nu} = \eta_{\epsilon\sigma}$$

$$\Lambda^\mu_\epsilon \simeq \delta^\mu_\epsilon + \omega^\mu_\epsilon$$

$$\Rightarrow (\delta^\mu_\epsilon + \omega^\mu_\epsilon) (\delta^\nu_\sigma + \omega^\nu_\sigma) \eta_{\mu\nu} = \eta_{\epsilon\sigma}$$

$$\Rightarrow \delta_\rho^\mu \omega^\nu_\sigma \eta_{\mu\nu} + \delta_\sigma^\nu \omega^\mu_\rho \eta_{\mu\nu} = 0$$

$$\omega_{\rho\sigma} + \omega_{\sigma\rho} = 0 \Rightarrow \omega_{\mu\nu} : \text{anti-symmetric}$$

6 independent elements

$$\phi'(x') = \phi(\Lambda^\mu_\nu x^\nu)$$

$$\delta_\omega \phi = \omega^\mu_\nu x^\nu \partial_\mu \phi(x)$$

$$\delta_\omega \mathcal{L} = \partial_\mu (\omega^\mu_\nu x^\nu \mathcal{L})$$

$$\Rightarrow j^\mu = -\eta^{\mu\nu} \partial_\nu \phi (\omega^\rho_\sigma x^\sigma) - (\omega^\mu_\nu x^\nu \mathcal{L})$$

$$= \omega_{\rho\sigma} (T^{\mu\rho} x^\sigma - T^{\mu\sigma} x^\rho)$$

$$= \omega_{\rho\sigma} M^{\mu[\rho\sigma]}$$

$$M^{\sigma[\rho\sigma]} \equiv M^{\rho\sigma} = T^{\sigma\rho} x^\rho - T^{\sigma\rho} x^\nu \Rightarrow \text{angular momentum}$$

- Global U(1)

$$\phi \rightarrow e^{i\alpha} \phi$$

$$\phi^* \rightarrow e^{-i\alpha} \phi^*$$

$$\Rightarrow \begin{aligned} \delta_\alpha \phi &= i\alpha \\ \delta_\alpha \phi^* &= -i\alpha \end{aligned} \quad \alpha : \text{constant}$$

$$\mathcal{L} = -(\partial_\mu \phi)(\partial^\mu \phi^*) + \pi^2 \phi \phi^* \Rightarrow \delta \mathcal{L} = 0$$

$$\frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} = -\partial^\mu \phi^* \quad \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi^*)} = -\partial^\mu \phi$$

$$\Rightarrow J^\mu = \alpha \left( -\phi \partial^\mu \phi^* + \phi^* \partial^\mu \phi \right) \rightsquigarrow \text{probability current}$$

"electric charge"

Question: What if  $\alpha$  became a function on the spacetime? :  $\alpha = \alpha(x)$

$\Rightarrow \mathcal{L}$  is not invariant.

$\Rightarrow$  Cure: Introduce a new field: Gauge field  $A_\mu$



# IV. Local Gauge Theory and Classical Electrodynamics

## IV.1. Recall: Electrodynamics

- Maxwell equations

$$\left\{ \begin{array}{l} \vec{\nabla} \cdot \vec{E} = 4\pi\rho \\ \vec{\nabla} \cdot \vec{B} = 0 \end{array} \right. \quad \left\{ \begin{array}{l} \vec{\nabla} \times \vec{B} - \frac{\partial \vec{E}}{\partial t} = 4\pi \vec{J} \\ \vec{\nabla} \times \vec{E} + \frac{\partial \vec{B}}{\partial t} = 0 \end{array} \right.$$

$(\vec{E}, \vec{B})$ : Electric / Magnetic field

$(\rho, \vec{J})$ : Electric density (current)

$$\Rightarrow \vec{E} = -\vec{\nabla}\phi - \frac{\partial \vec{A}}{\partial t}, \quad \vec{B} = \vec{\nabla} \times \vec{A}$$

Q. How can we describe the Maxwell equation in a covariant way?

$\Rightarrow$  Introduce four-potential  $A_\mu = (-\phi, \vec{A})$

anti-symmetric tensor  $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$

four-current  $J^\mu = (\rho, \vec{J})$

$$\Rightarrow \text{Maxwell eqn.} \Rightarrow \left\{ \begin{array}{l} \partial_\mu F^{\mu\nu} = 4\pi J^\nu \quad (\text{field eqn}) \\ \partial_\rho F_{\mu\nu} + \partial_\mu F_{\nu\rho} + \partial_\nu F_{\rho\mu} = 0 \quad (\text{Bianchi idet.}) \end{array} \right.$$

## IV.2. Local Gauge Theory in QM

Global  $U(1)$  gauge  $\psi \rightarrow e^{i\alpha} \psi$

$$\Downarrow \quad -\frac{1}{2m} \vec{\nabla}^2 (e^{i\alpha} \psi) = -\frac{1}{2m} \vec{\nabla}^2 \psi$$

Local  $U(1)$  gauge  $\psi \rightarrow e^{i\alpha(\vec{x})} \psi$

$$-\frac{1}{2m} \vec{\nabla}^2 (e^{i\alpha(\vec{x})} \psi) \neq -\frac{1}{2m} \vec{\nabla}^2 \psi$$

Q: How can we cure this problem?

A: Introduce new local function  $\vec{A}$  which transforms under local  $U(1)$  gauge as  $\left[ \begin{array}{l} \vec{A} \rightarrow \vec{A} + i\vec{\nabla}\alpha(x) \\ \phi \rightarrow \phi - \partial_t \alpha(x) \end{array} \right. \quad \psi \rightarrow \psi e^{ie\alpha(x)}$

Define kinematic momentum

$$\vec{\pi} = \vec{p} - e\vec{A} = -i(\vec{\nabla} - ie\vec{A})$$

$\uparrow$  canonical momentum

$\vec{D}$ : covariant derivative

$$H = \frac{1}{2m} \vec{\pi}^2 + e\phi$$

$$= \frac{1}{2m} (\vec{p} - e\vec{A})^2 + e\phi = -\frac{1}{2m} (\vec{\nabla}^2 - ie(\vec{\nabla} \cdot \vec{A} + \vec{A} \cdot \vec{\nabla}) - e^2 \vec{A}^2) + e\phi$$

$$H \rightarrow -\frac{1}{2m} (\vec{\nabla}^2 - ie(\vec{\nabla} \cdot \vec{A} + i\vec{\nabla}^2 \alpha + \vec{A} \cdot \vec{\nabla} + (\vec{\nabla} \alpha) \cdot \vec{\nabla}) - e^2 (\vec{A}^2 + i(\vec{A} \cdot (\vec{\nabla} \alpha) + (\vec{\nabla} \alpha) \cdot \vec{A}) - (\vec{\nabla} \alpha)^2) + e(\phi - \frac{\partial \alpha}{\partial t})$$

$$\vec{\nabla}^2 (e^{ie\alpha} \psi) = \vec{\nabla} \cdot (ie(\vec{\nabla}\alpha)\psi + \vec{\nabla}\psi) e^{ie\alpha}$$

$$\begin{aligned} H\psi \rightarrow & -\frac{1}{2m} e^{ie\alpha} \left[ (ie(\vec{\nabla}^2\alpha)\psi + ie(\vec{\nabla}\alpha) \cdot (\vec{\nabla}\psi) - e^2 (\vec{\nabla}\alpha)^2 \psi \right. \\ & \left. + (\vec{\nabla}^2\psi) + ie(\vec{\nabla}\psi) \cdot (\vec{\nabla}\alpha) \right) \\ & - ie(\vec{\nabla} \cdot \vec{A} + i\vec{\nabla}^2\alpha + (-ie\vec{A} \cdot (\vec{\nabla}\alpha)\psi + \vec{A} \cdot (\vec{\nabla}\psi)) \\ & + ie(\vec{\nabla}\alpha)^2 \psi + (\vec{\nabla}\alpha) \cdot (\vec{\nabla}\psi)) \\ & - e^2 (\vec{A}^2 \psi + i(\vec{A} \cdot (\vec{\nabla}\alpha) + (\vec{\nabla}\alpha) \cdot \vec{A}) \psi - (\vec{\nabla}\alpha)^2 \psi) \\ & \left. + e(\phi - \frac{\partial\alpha}{\partial t}) \psi e^{ie\alpha} \right] \end{aligned}$$

$\Rightarrow$  gauge invariant equation

$$H\psi = -i \frac{\partial}{\partial t} \psi \quad \text{is invariant under local } U(1)$$

Cost : Replace  $\vec{p} \rightarrow \vec{\pi}$   
Introduce new local field  $\vec{A}, \phi$   
gauge field

### IV.3. Lagrangian Description of Maxwell Theory

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + 4\pi A_\mu j^\mu \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$\text{Field equation} \quad \partial_\mu F^{\mu\nu} = 4\pi j^\nu \rightarrow \partial_\nu j^\nu = 0$$

$$\square A^\nu - \partial^\nu (\partial_\mu A^\mu) = 4\pi j^\nu$$

Note We can find electrodynamical Lagrangian

$$\mathcal{L} = \frac{1}{2} (\vec{E}^2 - \vec{B}^2)$$

• Gauge invariance

$$A^\mu \rightarrow A^\mu + \partial^\mu \alpha(x) \quad \rightarrow F_{\mu\nu} \rightarrow F_{\mu\nu}$$

Choose Lorenz gauge condition

$$\partial_\mu A^\mu = 0 \quad \rightarrow \quad \partial_\mu A'^\mu = \partial_\mu A^\mu + \partial_\mu (\partial^\mu \alpha) = 0$$

$$\Rightarrow \square \alpha = 0$$

$$\Rightarrow \square A^\mu = 4\pi j^\mu : \text{Klein-Gordon equation with current } j^\mu$$

#### IV.4. Scalar Electrodynamics

Local U(1) gauge invariant theory

Covariant  
derivative

$$\mathcal{L}_{\text{scalar}} = -\frac{1}{2} (D_\mu \phi) (D^\mu \phi)^* + \frac{1}{2} m^2 |\phi|^2 \quad D_\mu \phi = (\partial_\mu - igA_\mu) \phi$$

$$D_\mu \phi^* = (\partial_\mu + igA_\mu) \phi^*$$

$\mathcal{L}_{\text{scalar}}$  is invariant under  $\phi \rightarrow e^{ig\alpha(x)} \phi$ ,  $\phi \rightarrow e^{-ig\alpha(x)}$   
 $A_\mu \rightarrow A_\mu + \partial_\mu \alpha(x)$

$\Rightarrow$  Noether theorem

$$j^\mu = ig (\phi^* D^\mu \phi - \phi (D^\mu \phi)^*)$$

If we want to consider  $A_\mu$  as a dynamical field, we should add  $\mathcal{L}_{EM}$  to  $\mathcal{L}_{\text{scalar}}$

$$\mathcal{L} = \mathcal{L}_{\text{scalar}} + \mathcal{L}_{EM}$$

$$= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - (D_\mu \phi) (D^\mu \phi)^* + \frac{1}{2} m^2 |\phi|^2$$

$$= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - (\partial_\mu \phi) (\partial^\mu \phi) + \frac{1}{2} m^2 |\phi|^2$$

$$+ ig A_\mu (\phi^* \partial^\mu \phi - \phi \partial^\mu \phi^*) - g^2 A^2 |\phi|^2$$

# V. Spontaneous Symmetry Breaking

## V.1. Prelude: Ising model and Landau theory

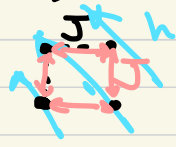
(Ising, 1924) Description of (ferro  $\rightarrow$ ) magnetism by spin operator

$$H(s_i; \{K\}) = - \sum_{\langle i, j \rangle} K^{(n)} G_{i(j)}^{(n)} \quad \begin{matrix} G: \text{local operator} \\ K: \text{coupling constant} \end{matrix}$$

- ex)  $n=1 \quad G_i^{(1)} = s_i$
- $n=2 \quad G_{ij}^{(2)} = s_i s_j$
- $n=3 \quad G_{ijk}^{(3)} = s_i s_j s_k \dots$

We only consider  $n=1$  and  $n=2$  (up to nearest neighbour)

$$H(s; J, h, T) = -h \sum_{i=1}^N s_i - \sum_{\langle i, j \rangle} J_{ij} s_i s_j$$



\*  $h$  breaks time-reversal symmetry ( $s \rightarrow -s$ ) Zeeman exchange interaction  
 $J > 0$  Ferromagnetic Ising   
 $J < 0$  Anti - "

Suppose  $J=0$  (paramagnetic case)

$$Z[s; J=0, h, T] = \prod_{i=1}^N (e^{h\beta} + e^{-h\beta}) = (2 \cosh^N h\beta)$$

$$m = -\frac{1}{N} \frac{\partial F}{\partial h} = \tanh h\beta$$

When  $J \neq 0$ , each spin feels not only the external field  $h$ , but also an effective interaction  $h_{eff}$  due to the all other spins.

But, in general,  $h_{eff}$  is unknown.

$$H = - \sum_{\langle i, j \rangle} J_{ij} s_i s_j - h \sum s_i$$

$$= - \sum_i s_i \left( h + \sum_j J_{ij} s_j \right) \equiv - \sum_i h_i^{eff} s_i$$

$$h_i^{eff} = h + \sum_j J_{ij} s_j = h + \sum_j J_{ij} \langle s_j \rangle + \sum_j J_{ij} (s_j - \langle s_j \rangle)$$

order parameter

$$= h + 2d J m$$

Assume  
No fluctuation on  $s_j \Rightarrow$  MFT

$$\therefore m = \tanh(\beta h_{eff}) = \tanh \beta (h + 2d J m)$$

- Landau theory  
Assumptions

$$h, J, t = \frac{T - T_c}{T_c}$$

Let  $\mathcal{F}_L$ : energy functional with respect to the  $\{K\}$  and order parameter  $\phi \rightarrow m$

i)  $f_L$  obeys symmetries of the system.

ii) Near critical temperature  $T_c$ ,  $f_L$  is an analytic function for both  $\xi k$  and  $\phi$

$$f_L = \sum a_n (\xi k) \phi^n$$

• Landau theory for Ising model

If  $h=0 \Rightarrow$  time reversal symmetry exists

$\Rightarrow$  We demand  $f_L[\phi] = f_L[-\phi]$

$$\Rightarrow f_L = a_0 + a_2 \phi^2 + a_4 \phi^4 + \dots$$

• We set  $a_0 = 0$  since it is just an energy shift.

•  $a_4$  must be positive.

$$\Rightarrow f_L = a \phi^2 + b \phi^4$$

At the minimum:  $\left. \frac{\partial f_L}{\partial \phi} \right|_{\phi=\phi_*} = 0 \Rightarrow \phi_* = 0$  or  $\phi_* = \pm \sqrt{\frac{-a}{2b}}$

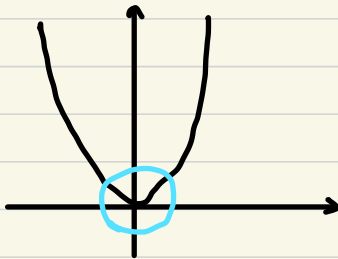
If  $a(t) \sim a_0 t + a_1 t^2 + \mathcal{O}(t^3)$

Since  $\phi(t=0) = 0 \Rightarrow a_0 = 0$

$$\Rightarrow f_L = a t \phi^2 + b \phi^4$$

①  $t > 0 \Rightarrow$  Global minimum occurs at  $\phi_* = 0$

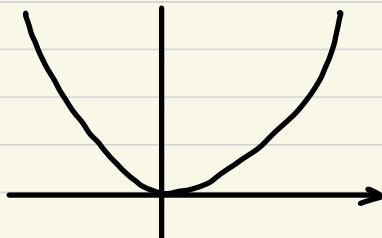
$\downarrow$   
 $T > T_c$



$$f_L \sim \phi^2$$

②  $t = 0$

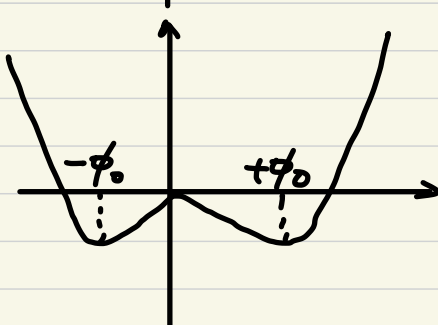
$\downarrow$   
 $T = T_c$



$$f_L \sim \phi^4$$

③  $t < 0$

$\downarrow$   
 $T < T_c$



$$\phi_* = 0 \text{ or } \pm \sqrt{\frac{-a t}{2b}} \equiv \pm \phi_0$$

•  $\phi = 0$  is not only no longer the global minimum, but also the local maximum.

• There exist two minimum related to the symmetry  $\phi \rightarrow -\phi$   
 $\Rightarrow$  Spontaneous symmetry breaking

## V.2. Spontaneous Symmetry Breaking

- Double-well potential

• Self-interacting field

$$\mathcal{L} = -\frac{1}{2} (\partial_\mu \phi)^2 + \underbrace{\frac{1}{2} m^2 \phi^2}_{\text{self-interacting term}} + \underbrace{V(\phi)}_{\frac{a}{3} \phi^3 + \frac{b}{4} \phi^4 + \dots}$$

Note Coefficient of  $\phi^2$  term is called mass of field.

$$\Rightarrow \text{Equation of motion: } (\square - m^2) \phi = V'(\phi)$$

• Mexican hat potential

$$\mathcal{L}_0 = -\frac{1}{2} (\partial_\mu \phi) (\partial^\mu \phi) - \underbrace{\left( -\frac{a^2}{2} \phi^2 + \frac{b^2}{4} \phi^4 \right)}_{V(\phi)}$$

• Symmetry

- \* Translation ] Poincaré
- \* Lorentz
- \*  $\phi \rightarrow -\phi$  -  $\mathbb{Z}_2$

$$V'(\phi) = 0 \Rightarrow \phi = 0 \text{ or } \pm \frac{a}{b}$$

Both Poincaré and  $\mathbb{Z}_2$  invariant

Poincaré invariant  
 $\mathbb{Z}_2$  - not invariant

- Spontaneous

Noether current: Energy functional

$$E = \int_V dV \left( \frac{1}{2} (\partial_0 \phi)^2 + \frac{1}{2} (\vec{\nabla} \phi)^2 - \frac{a^2}{2} \phi^2 + \frac{b^2}{4} \phi^4 \right)$$

$$\delta E = \int_V dV \left\{ (\partial_0 \phi) (\partial_0 \delta \phi) + (\vec{\nabla} \phi) (\vec{\nabla} \delta \phi) - a^2 \phi \delta \phi + b^2 \phi^3 \delta \phi \right\}$$

$$= \int_V dV \delta \phi \left[ -(\partial_0 \phi)^2 - (\vec{\nabla} \phi)^2 - a^2 \phi + b^2 \phi^3 \right]$$

$$\Rightarrow \phi = 0 \text{ or } \pm \frac{a}{b} : \text{critical point}$$

• local maximum    • global minimum

• unstable point    • stable point

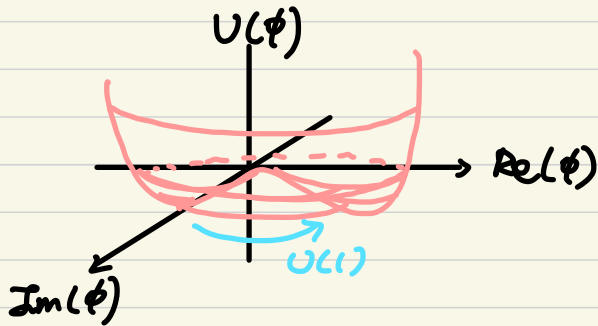
$\Rightarrow$  ground state of classical field

$\Rightarrow$  In this case, ground state is less symmetric ~~Poin  $\times$   $\mathbb{Z}_2$~~   $\rightarrow$  Poin  
\* If  $a^2 \rightarrow -a^2$ , there is no SSB.

- SSB for Complex field

$$\mathcal{L} = -\frac{1}{2}(\partial_\mu \phi)(\partial^\mu \phi)^* - \underbrace{\left(-\frac{a^2}{2}|\phi|^2 + \frac{b^2}{4}|\phi|^4\right)}_{\text{assume } a, b > 0}$$

• Symmetry: Poin  $\times$   $U(1)$        $\phi \rightarrow e^{i\alpha} \phi$



$$V'(\phi) = 0 \Rightarrow \frac{b^2}{4} \varphi |\varphi|^2 - \frac{1}{2} a^2 \varphi = 0$$

$$\Rightarrow \begin{cases} |\varphi| = \frac{a}{b} & \Rightarrow \varphi = \frac{a}{b} e^{i\theta} \leadsto \text{topologically circle} \\ \varphi = 0 & \leadsto \varphi \sim \varphi e^{i\alpha + \theta} \end{cases}$$

• Energy functional

$$E = \int d^3\vec{x} \left( \frac{1}{2} |\partial_0 \phi|^2 + \frac{1}{2} |\vec{\nabla} \phi|^2 - \frac{a^2}{2} |\phi|^2 + \frac{b^2}{4} |\phi|^4 \right)$$

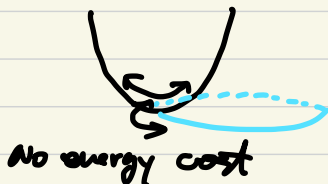
Let  $\phi = \rho(x) e^{i\omega(x)}$

$$E = \int d^3\vec{x} \left[ \frac{1}{2} (\partial_0 \rho)^2 + \frac{1}{2} (\vec{\nabla} \rho)^2 + \frac{\rho^2}{2} ((\partial_0 \omega)^2 + (\vec{\nabla} \omega)^2) - \frac{a^2}{2} \rho^2 + \frac{b^2}{4} \rho^4 \right]$$

Local variation

$$\delta E = -\frac{1}{4} \frac{a^4}{b^2} + \int d^3\vec{x} \left[ \frac{1}{2} (\delta \partial_0 \rho)^2 + \frac{1}{2} (\delta \vec{\nabla} \rho)^2 + \frac{1}{2} \left(\frac{a}{b}\right)^2 ((\delta \partial_0 \omega)^2 + (\delta \vec{\nabla} \omega)^2) + a^2 (\delta \rho)^2 \right]$$

$\Rightarrow \delta \rho = 0, \delta \omega = \text{const} \Rightarrow$  No energy change  
Otherwise, energy is increased.



### V.3. Goldstone Theorem

- Linearisation

Extremum

:  $\phi_0$

Near extremum point (stable)  $\phi(x) = \phi_0(x) + \varphi(x)$   
 $\hookrightarrow$  satisfying eom

Ex) - Gravity theory  $g_{\mu\nu}(x) = \eta_{\mu\nu} + h_{\mu\nu}(x)$   
 $\hookrightarrow$  Graviton in flat space  
 $G_{\mu\nu}[g] = \underbrace{G_{\mu\nu}[\eta]}_0 + \underbrace{G_{\mu\nu}[\eta, h]}_0$

•  $\phi^4$  - theory of real scalar

$$\mathcal{L} = -\frac{1}{2} (\partial_\mu \phi)^2 + \frac{m^2}{2} \phi^2 + \frac{\lambda}{4} \phi^4$$

EOM  $\square \phi + m^2 \phi + \lambda \phi^3 = 0$

Non-linear PDE

$$\phi = \phi_0 + \varphi$$

$$\square \phi_0 + \square \varphi + m^2 \phi_0 + m^2 \varphi + \lambda (\phi_0 + \varphi)^3 = 0$$

$$\Rightarrow \square \varphi + m^2 \varphi + \lambda (3\phi_0 \varphi^2 + 3\phi_0^2 \varphi + \varphi^3) = 0$$

assuming small approximation

$$\Rightarrow \square \varphi + (m^2 + 3\lambda \phi_0^2) \varphi = 0 \quad \text{Linear PDE}$$

$\hookrightarrow \varphi$  is a scalar field with eff. mass  $m^2 + 3\lambda \phi_0^2 \equiv m_{\text{eff}}^2$

In formal case  $m^2 = -a^2, \lambda \rightarrow b^2$

$$\phi_0 = \pm a/b \Rightarrow m_{\text{eff}}^2 = 2a^2 > 0 \Rightarrow \text{stable}$$

$$\phi_0 = 0 \Rightarrow m_{\text{eff}}^2 = -a^2 < 0 \Rightarrow \text{unstable}$$

- Goldstone Theorem

Complex scalar theory  $\phi = \rho(x) e^{i\omega(x)}$

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \rho \partial^\mu \rho - \frac{1}{2} \rho^2 \partial_\mu \omega \partial^\mu \omega - \frac{1}{2} a^2 \rho^2 + \frac{1}{4} b^2 \rho^4$$

$$\Rightarrow \delta_\rho \mathcal{L} = \square \rho - \rho \partial_\mu \omega \partial^\mu \omega - a^2 \rho + b^2 \rho^3 = 0$$

$$\delta_\omega \mathcal{L} = \partial_\mu (\rho^2 \partial^\mu \omega) = 0 \Rightarrow \text{conservation law}$$

$$\rho_0 = \frac{a}{b}, \quad \omega_0 = \text{const}$$

$$\Rightarrow \rho = \frac{a}{b} + \varphi \quad \omega = \omega_0 + \theta$$

$$\mathcal{L} = -\frac{1}{2} (\partial_\mu \varphi) (\partial^\mu \varphi) - \frac{1}{2} \left(\frac{a}{b}\right)^2 (\partial_\mu \theta) (\partial^\mu \theta) - \frac{1}{2} \frac{a^4}{b^2} + a^2 \varphi^2 + (\text{Higher order})$$



com:  $\begin{cases} \square \varphi - 2a^2 \varphi = 0 & \Rightarrow \text{massive field} \\ \square \theta = 0 & \Rightarrow \text{massless field} \end{cases}$

Goldstone theorem: Broken symmetry  $\Rightarrow$  Massless field  
 $\Rightarrow$  Goldstone field

#### V.4. Abelian Higgs Model

$$\begin{aligned} \mathcal{L} &= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - (D_\mu \phi) (D^\mu \phi)^* - V(\phi) \\ &= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - (\partial_\mu \phi) (\partial^\mu \phi)^* - ig A_\mu (\phi \partial^\mu \phi^* - \phi^* \partial^\mu \phi) \\ &\quad - g^2 A^2 |\phi|^2 - \frac{a^2}{2} |\phi|^2 - \frac{b^2}{4} |\phi|^4 \\ \phi &= \rho(x) e^{i\omega(x)} \end{aligned}$$

$$\begin{aligned} \mathcal{L} &= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \partial_\mu \rho \partial^\mu \rho - \rho^2 \partial_\mu \omega \partial^\mu \omega \\ &\quad - ig A_\mu (\rho(x) e^{i\omega} [\cancel{\partial^\mu \rho} e^{-i\omega} + \rho (-i \partial^\mu \omega) e^{-i\omega}] \\ &\quad - \rho(x) e^{-i\omega} [\cancel{\partial^\mu \rho} e^{i\omega} + \rho (+i \partial^\mu \omega) e^{i\omega}]) \\ &\quad - g^2 A^2 \rho^2 - \frac{a^2}{2} \rho^2 - \frac{b^2}{4} \rho^4 \end{aligned}$$

$$\begin{aligned} &= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \partial_\mu \rho \partial^\mu \rho - \rho^2 \partial_\mu \omega \partial^\mu \omega \\ &\quad - 2g A_\mu \rho^2 \partial^\mu \omega - g^2 A^2 \rho^2 - \frac{a^2}{2} \rho^2 - \frac{b^2}{4} \rho^4 \end{aligned}$$

Ground state  $\rho(x) = \frac{a}{b}$ ,  $\omega(x) = \omega_0$ ,  $A_\mu = 0$   
(Poincaré invariant)

Linearisation:  $\rho(x) = \frac{a}{b} + H(x)$ ,  $\omega(x) = \omega_0 + \pi(x)$ ,  $A_\mu = 0 + Q_\mu(x)$

Introduce a new field  $B_\mu = Q_\mu + \frac{1}{g} \partial_\mu \pi$

$$\mathcal{L} \approx -\frac{1}{4} F_{\mu\nu}^B F^{B\mu\nu} - \left(\frac{ag}{b}\right)^2 B_\mu B^\mu - \frac{a^2}{2} H^2 - \partial_\mu H \partial^\mu H$$

Spin-1 Lagrangian with mass  $\frac{ag}{b} \equiv \mu$

Scalar field Lagrangian

$\Rightarrow$  Proca Lagrangian

• eqn for  $B_\mu$   $\left[ \begin{array}{l} (\square - \mu^2) B_\mu = 0 \\ \partial^\mu B_\mu = 0 \end{array} \right] \Rightarrow$  3 independent component (scalar dof)

- Massive vector  $\Rightarrow$  No gauge invariance
- Mass of B is related to charge g and mass of H.

⇒ Near ground state : Massless vector  
Scalar



$SU(2) \times U(1)$  : Electroweak



Massive vector :  $B_\mu$   
Massive scalar :  $H$

Higgs phenomena

Higgs

$$a=1,2,3,0 \quad B_\mu^a \begin{cases} A \\ W^\pm \\ Z \end{cases}$$

